

nnis ACS (WETLabs) Protocols - Underway Flowthru D

Collection

- Non-Filtered ACS, Filtered(Cole Palmer 0.2 micron filter) A
 — Thermosalinograph, Fluorometer and Backscattering (EC
- Thermosalinograph necessary for temperature and salinit of AC measurements
- Water pumped from ~3m depth
- 3 Nanopure Water Calibrations (1 before and 2 during cru
- DH4 used to log all flowthru instruments (hourly files) and real-time using Wetview
- WAP used to time merge all flowthru data



- Interpolate ACS (Filtered and non-Filtered) absorption and beam-c data to same hyperspectral channel set (c1)
- Temperature and Salinity correction of ACS absorption and beam-c using coincident ship flowthru Thermosalinograph data
- Temperature correct nanopure water calibration file
- Subtract the nanopure water calibration from the ACS data (absorption and beam-c)
- Remove spikes normally caused by bubbles in system using a standard deviation filter (x4) – all records removed
- Compute Particulate Absorption: $a_p = a_t$ (non-filtered ACS) a_q (filtered ACS)
- Apply scatter correction (Rottgers 2013) to a_p
- Add scatter corrected particulate absorption (a_p) to CDOM absorption (a_g) yielding total absorption without water (a_{t-w})
- Add spectral pure water absorption coefficients (a_w) to total absorption without water (a_{t-w}) yielding total absorption (a_t).
- Particulate absorption (a_p), spectral scattering (b), omega (b/c) can be computed/created in spreadsheet (not written to corrected/processed files).

s ACS (WETLabs) Protocols - Underway Flowthru Data

Cruise Flowthru Objectives

- Characterize the spatial variability of the water's optical properties (a,b->bb,c) along the cruise track and how the variability impacts the uncertainty of insitu measurements at each station and ocean color response.
- Determine the water total and dissolved absorption (at, ag) properties at specific wavelengths
- Define frontal boundaries (Gulf Stream and Coastal Fresh Water/Upwelled Areas) using Thermal, Biological and Optical Properties
- Validate VIIRS Temperature, Salinity, Chlorophyll and IOP (QAA/LMI) products.

Define ocean processes and water mass types

Validate the spatial variability with a VIIRS 750m pixel by binning (mean &



Figure - IOP setup shows water bath setup using the acs, ac9 instruments. The ac9,acs are located inside the PVC containers and were in a controlled/constant temperature bath during operation.

Stennis ASD Protocols

Setup (FLDSPEC Software) w/ Protocol

- Foreoptic / FOV = 10deg, #scans + dark currents each set (water, sky, reference plaque) = 5, Optimize before each set, Reference = 10 inch 10% reflectance gray card, #stations = 18, #station collects = 26,
- NRL, NOAA, USF ASDs set up and operated by S. Ladner (NRL)
- NRL, USM, NOAA, USF, CCNY was collected using same protocols and using NRL reference plaque. Groups also collected a set using their own reference plaque and maybe protocol?

Collection Sequence

- Optimize gray reference plaque then collect 5 spectra
- Optimize water then collect 5 spectra and DCs
- Optimize sun then collect 5 spectra and DCs
- Optimization changes integration time based on relation brightness of target
- Optical zenith angles for reference, water, sky = 135,
- Relative sensor azimuth angle to the sun = >90 deg avoiding shading/shadows from ship, surface contanglint, foam, etc.)
- Integration times ranged from 68 to 4352ms
- All groups/measurements were taken from Bow or Mic 01 deck and on starboard side unless sun/ship orient forced collection to port side
- All groups shot same grey reference plaque (NRL) and from same

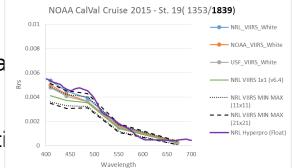


Figure: Ryan Vandermeulen (NASA) and Robert Arnone (USM) aboard the NOAA Ship Nancy Foster, taking ASD measurements reference plaque

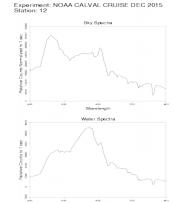
Stennis ASD Protocols (Cont)

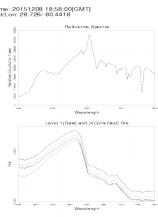
Processing

- Grey reference plaque has a know BRDF for normalizing the un-calibrated radiance measurement for Es
- Deriving Rrs from above water radiometry follow optics protocols from Mueller 2003 (Chapter 3 Method 2) $s = \frac{n}{i=0} \frac{CI_{NI}I_i}{n}$
- Sensor response signal (S) / reflectance is obtained from averaging N (5) readings from each of 3 targets and normalized to 1sec integration time $E_L \otimes_{E_c} \mathbb{F}_{sky} \square E_s = \frac{\mathbb{F}_L \mathbb{S}_g}{R_g} \mathbb{C} = \text{corrected DC}$ from Instrument, I_i is the integration time used for that reading, I_N is the normalized (1 second) integration time, N is the reading number.
- Water-Leaving Radiance (L_w) and Incident Spectral Irradiance (Irradiance (L_w) and Incident Spectral Irradiance
 - , F_L = unknown instrument radiance response calibration which cancel out in ratio for computing Rrs, R_q = gray reference plaque's bi-directional reflectance function (albedo)
- Rrs can be computed from the un-calibrated instrument data using p = Fresnel Reflectance for sky correction

expressed as the ratio of the radiance (or net signal) for the

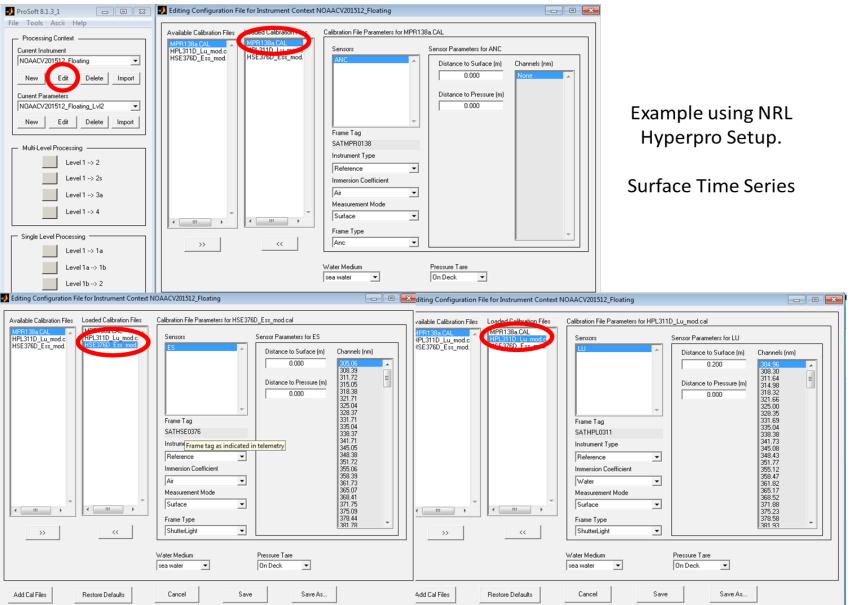
- The preferred Rrs correction (white) assumes from Carder and Steward 1985 that Rrs at 750nm should be black / zero. If not zero, then it is assumed that the error in the reflected skylight term ($S_{\rm sky}$) was not estimated correctly and is white (not wavelength
 - dependent) and may be subtracted from entire spengative reflectances in the NIR channels. Therefo smallest Rrs in the range from 700 to 825nm,
- For comparison to SNPP VIIRS the hyperspectral AS VIIRS Relative Spectral Response (RSR) table for each This RSR table (VIIRSN_IDPSv3_RSRs.txt) was ext NRL's APS software (based on NASA's SeaDAS Loriginated from the IDPS.
- <u>Blue Tile:</u> Slight modifications were made to the education derive the relative reflectance of the blue tile (R_{til}) blue tile measurements, the derived reflectance is simply





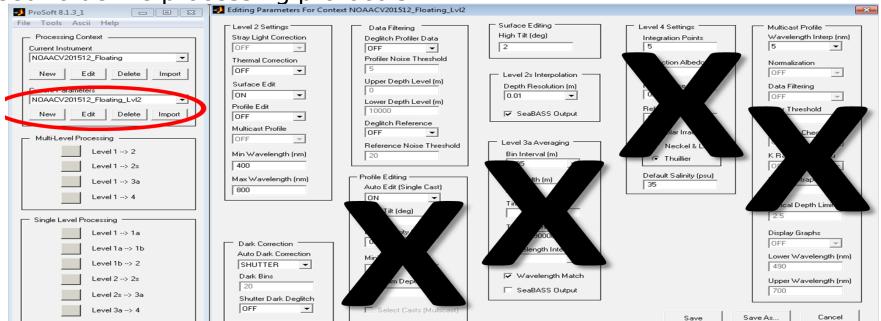
Stennis Hyperpro (Satlantic) Protocols

leed to define processing protocols



Stennis Hyperpro (Satlantic) Protocols (Cont)

leed to define processing protocols



X = Not used for floating package for this version (1.8.3) - Only level2 output worked yielding calibrated radiances.

all temporal samples.

Prosoft v8.1.4 (Notes from Prosoft Support Team / Darrell Adams) - New version w/ bugs fixed:

Level 3a processing of this data is possible if the Level 3a processing parameters are set appropriately. Time Interval and Time Width should not be set to 0 and 9000 respectively. For example when the Time Interval is set to 2 seconds and the Time Width is set to 1 second there are Level 3a data that is written to HDF and ASCII output files for each optical sensor (ie. Es and Ls). To output Lu, Ed, and Es at Levels 2s and 3a when all sensors are configured to be a "Reference" instrument type is possible if the Measurement Mode for all three is set to "Logger" in the Instrument Context. To have the Ed and Es (Irradiance optical sensor) wavelengths matched up to the Lu (radiance optical sensor) wavelengths the Level

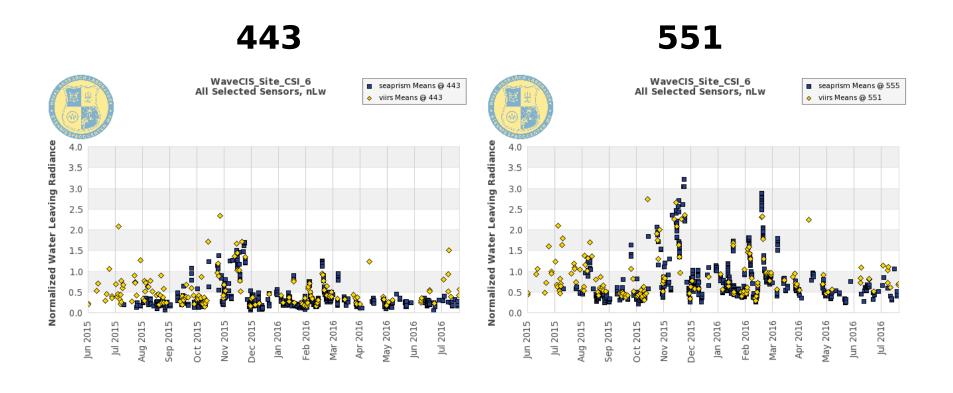
Stennis Hyperpro (Satlantic) Protocols (Cont)

Collection/Processing Rrs

- Floating HyperPro is a hyperspectral profiling radiometer that sim measures above-water downwelling irradiance (Ed) and in-water radiance (Lu) on a fixed floating platform and downwelling Es on
- Used to measure the hyper-spectral normalized water leaving rad computed using Thuiller hyperspectral extraterrestrial solar irrad Rrs = nLw(lamda)/F0(lamda)
- Both hyperspectral nLw and F0 were convolved to SNPP VIIRS RSF 2 Floating Hyperpro' (VIIRSN IDPSv3 RSRs.txt) obtained from NRL APS <- NASA SeaD
- Collected measurements at 21 stations.
- The spectral range of both Esd and Lu sensors is from 350 nm to 800 nm with 10 nm \pm 0.3 nm resolution.
- These instruments were used with a molded floatation collar, enabling in-water surface measurements to be taken over time just below the sea surface.
- The downwelling Ed sensor uses a cosine collector and is approximately 30 cm above the water surface which is different from ASD sky measurement.
- The upwelling (Lu) radiance sensor is mounted approximately 30 cm below the water surface.
- The ship mounted Es sensors was an cosine collector was mounted on the 01 deck with the
 other Hyperpros and was used for the downwelling computation of Rrs to omit uncertainty
 in the Ed (on the floating instrument) due to tilting caused from sea state and tension on
 the line.
- The USM Hyperpro was calibrated by NOAA- STAR in Oct 2015. The NRL Hyperpro was calibrated by NOAA STAR in Dec 2015. Calibration was similar to the Saclantic values and no major change was noted.
- The Floating HyperPro was deployed over the stern of the vessel. The tether was let out a sufficient distance from the boat (20 m to 30 m), allowing the instrument to float away from the boat to omit contamination from vessel-generated bubbles and shadowing or any other potential disturbances.

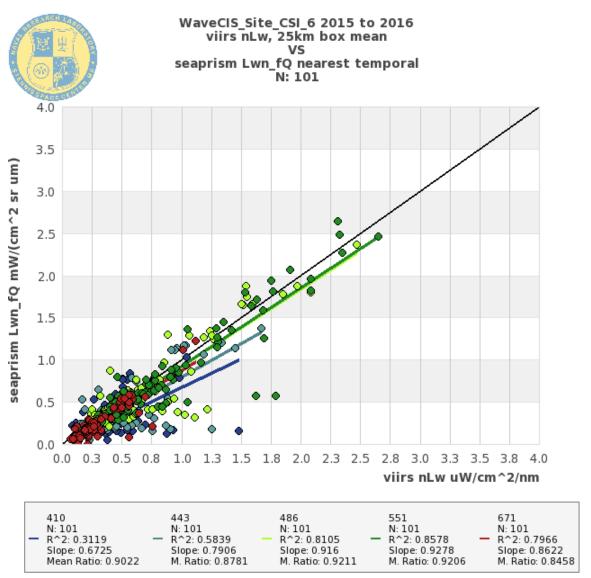
Mossurements were made over 5-10 minute time period

WaveCIS nLw FULL Time Series: NAVY APS 6.4 July 2015 - July 2016

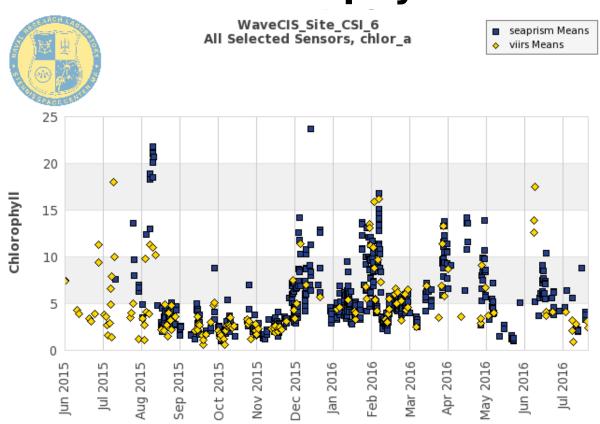


These are constrained results: 163 for VIIRS and 692 for WaveCIS (4-6 readings/day)

WaveCIS nLw Matchups (NAVY APS) July 2015 - July 2016



WaveCIS FULL Time Series: NAVY APS 6.4 Chlorophyll

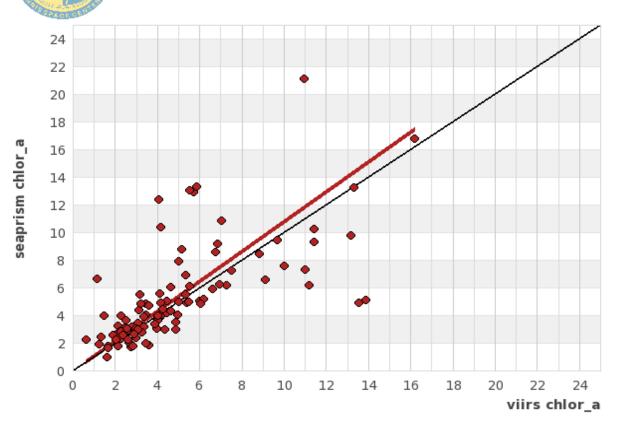


These are constrained results: 163 for VIIRS and 692 for WaveCIS (4-6 readings/day)

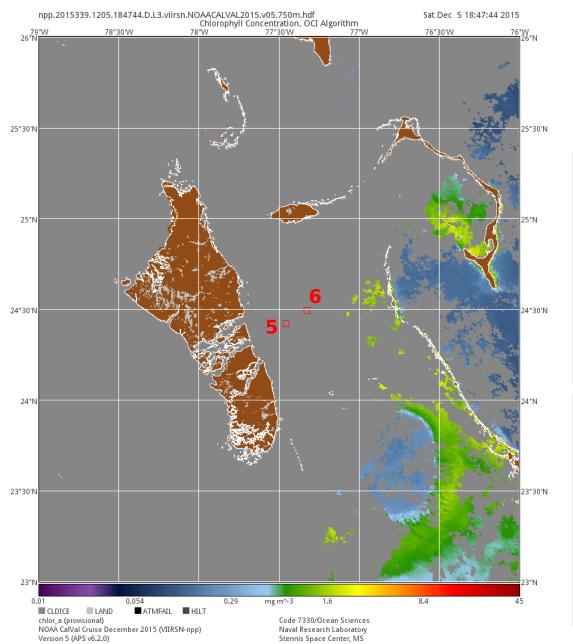
WaveCIS Chlorophyll Matchups (NAVY APS)

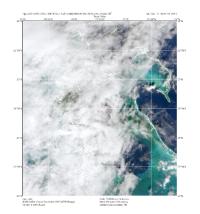
July 2015 - July 2016

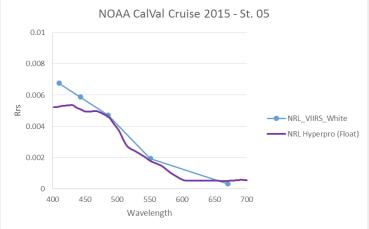
WaveCIS_Site_CSI_6 2015 to 2016
viirs chlor_a, 25km box mean
VS
seaprism chlor_a nearest temporal
N: 101

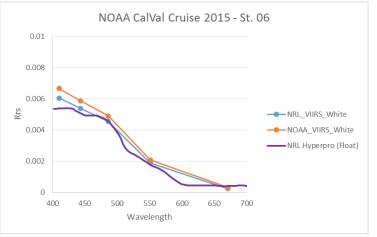


N: 101 R^2: 0.3672 Slope: 1.0772 Mean Ratio: 1.2387 12/05/16 (1847 GMT) St. 5,6

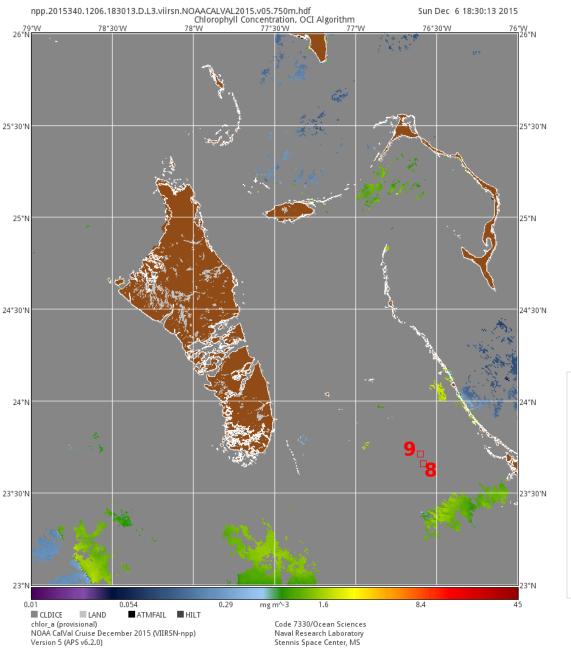


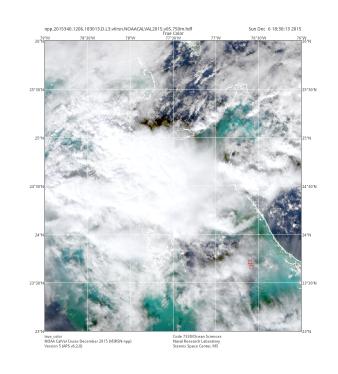


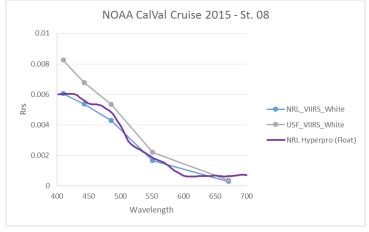




12/06/16 (1830 GMT) St. 8,9

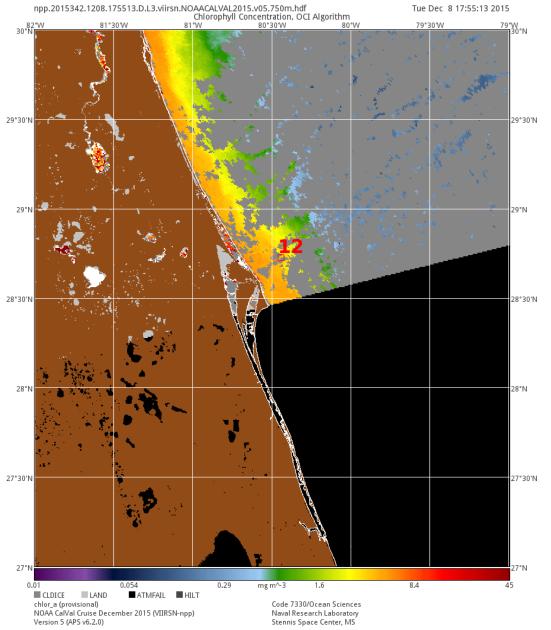


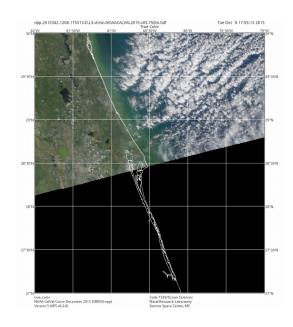


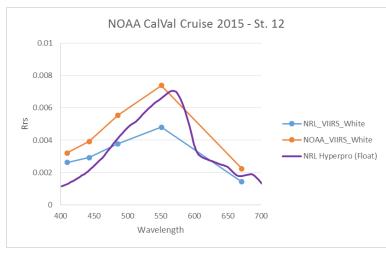


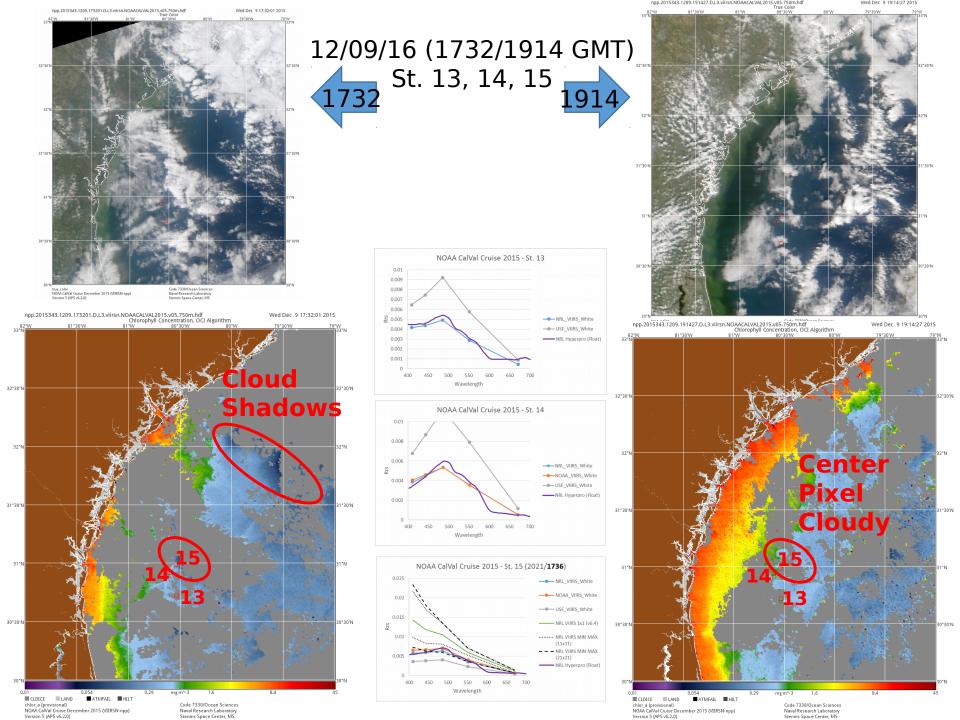
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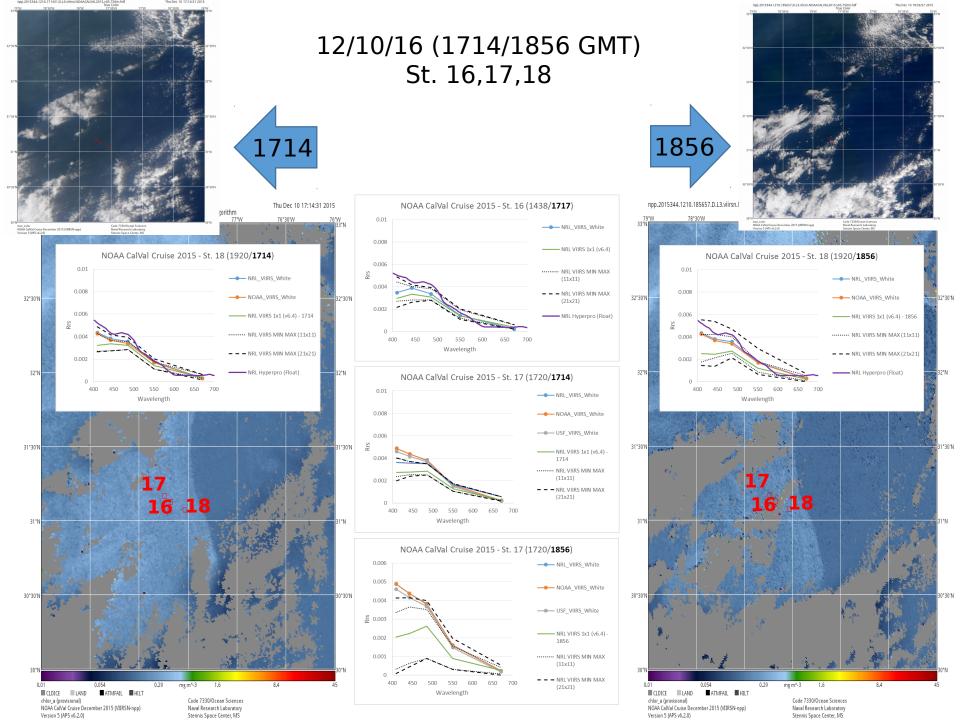


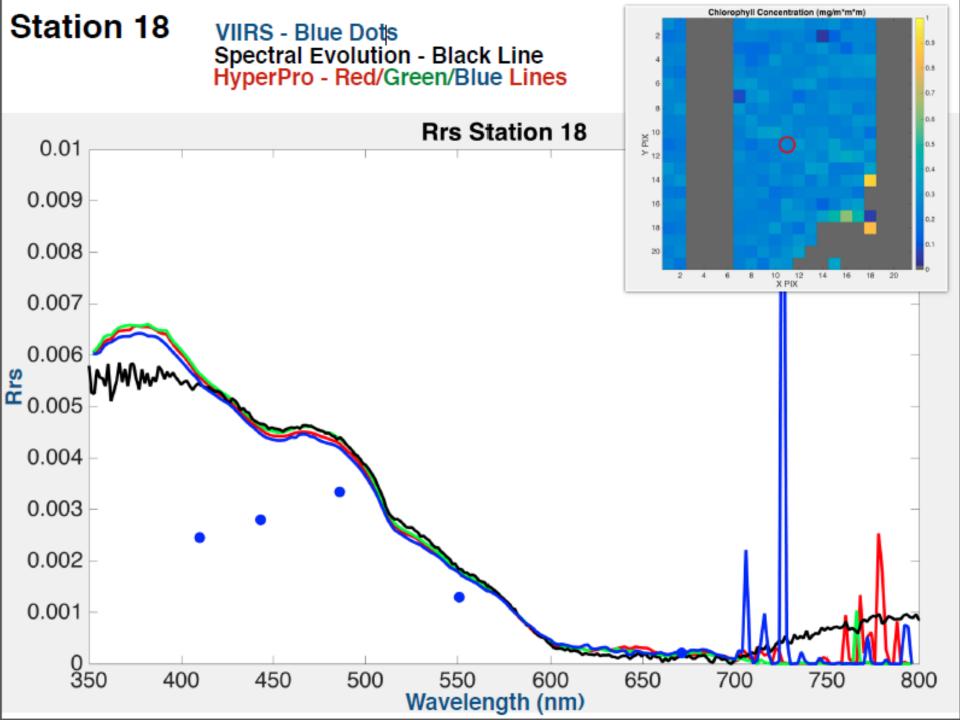








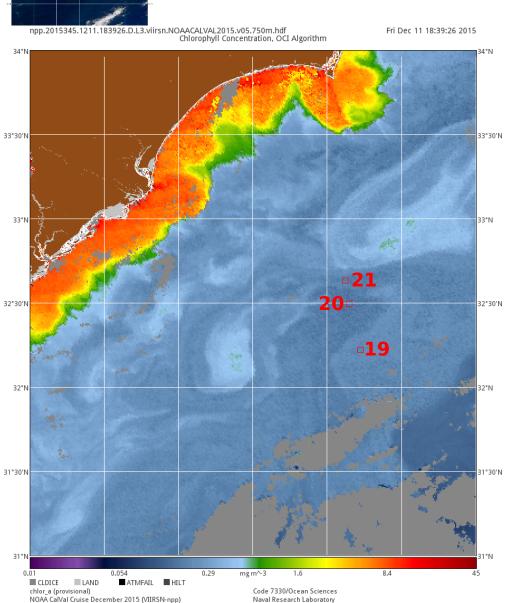




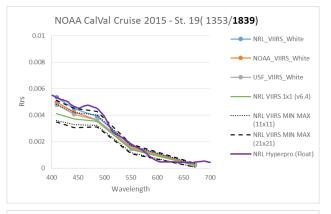


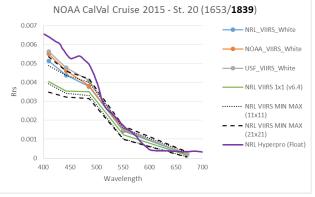
Version 5 (APS v6.2.0)

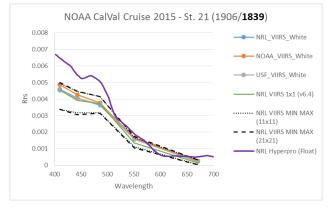
12/11/16 (1839 GMT) St. 19,20,21

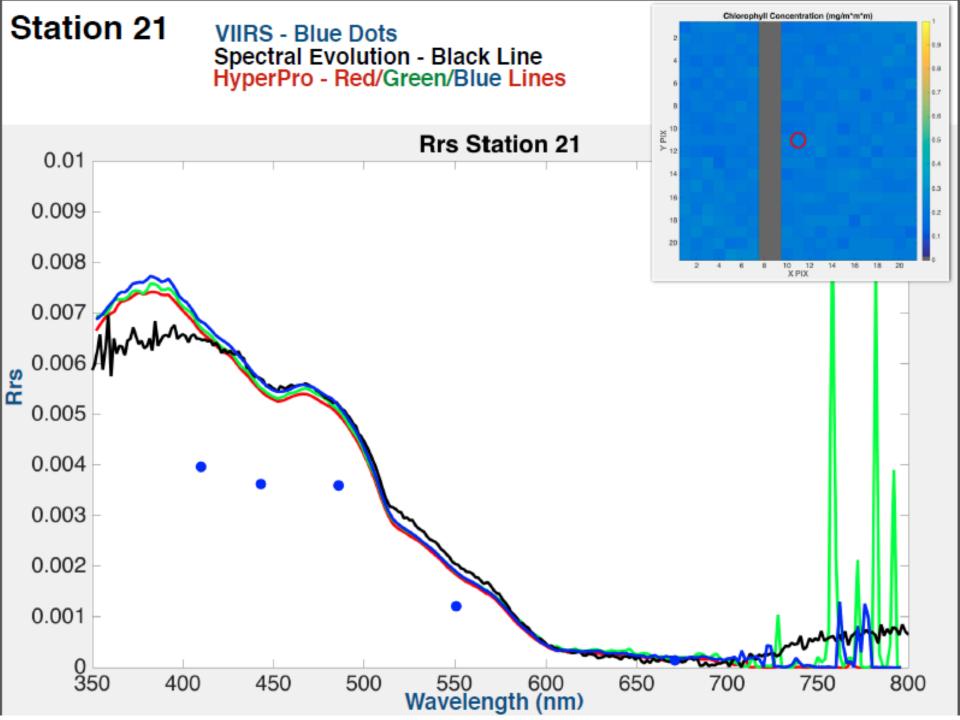


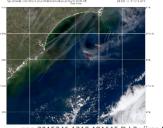
Stennis Space Center, MS







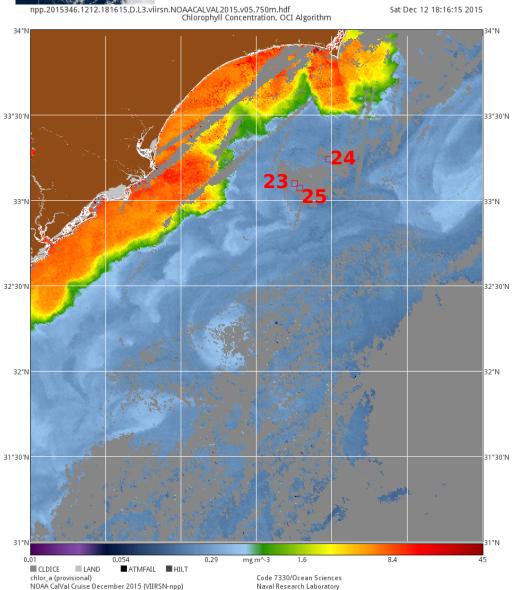




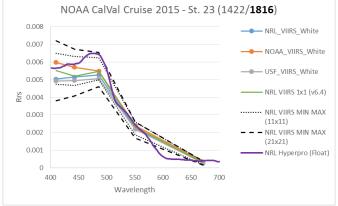
Version 5 (APS v6.2.0)

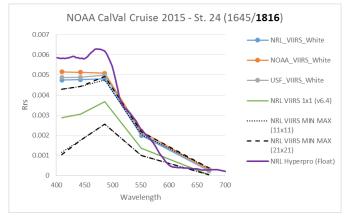
12/12/16 (1816 GMT)

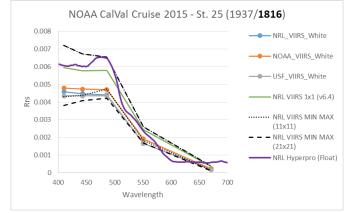
St. 23,24,25

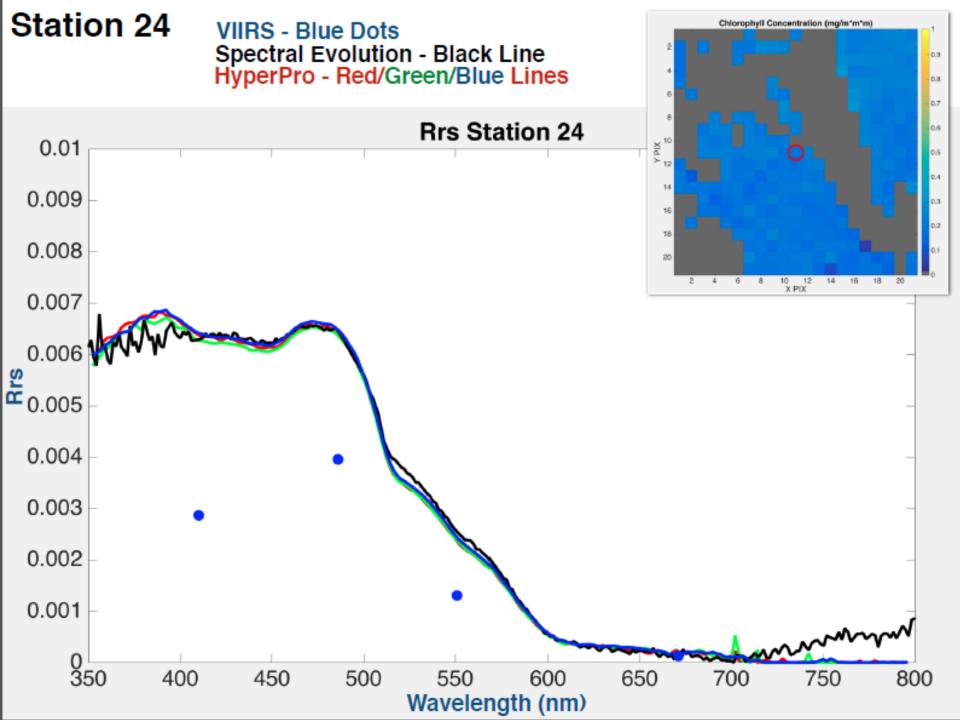


Stennis Space Center, MS

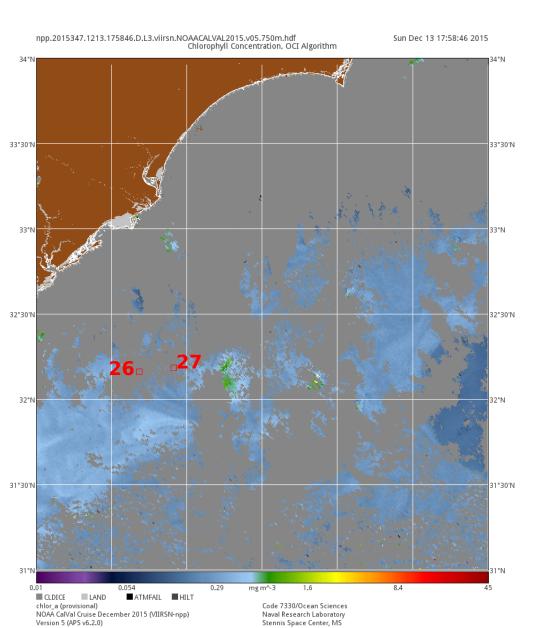


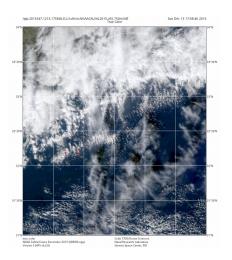


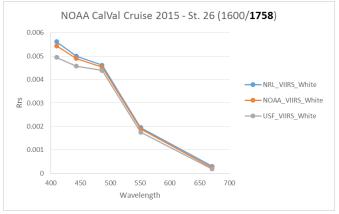


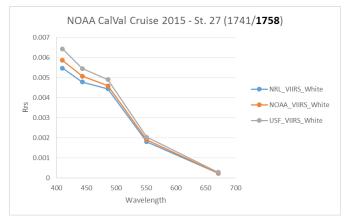


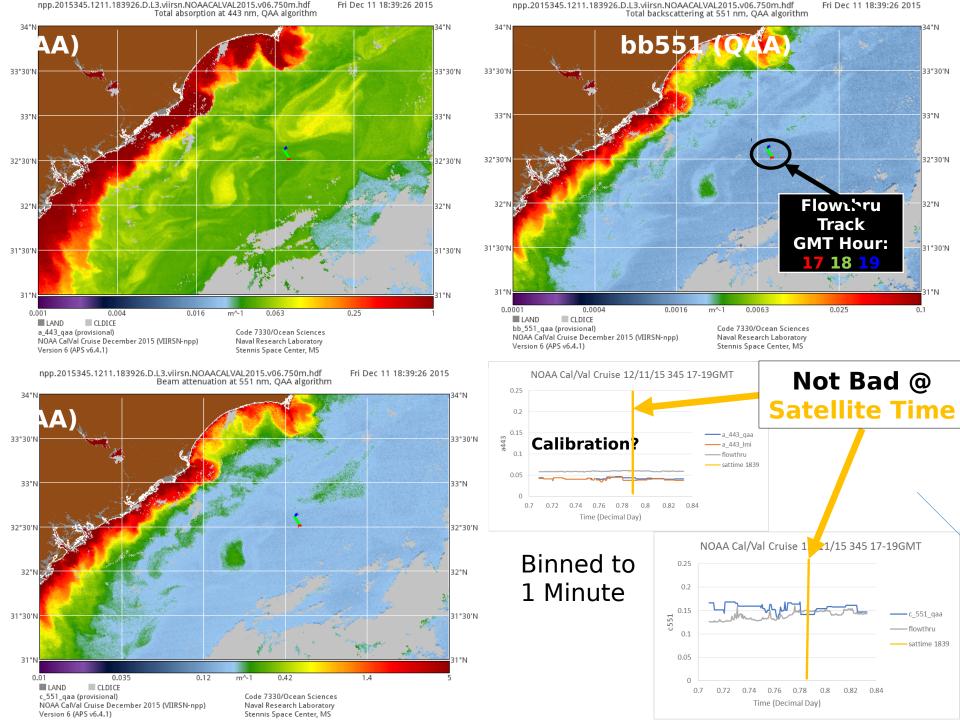
12/13/16 (1758 GMT) St. 26,27



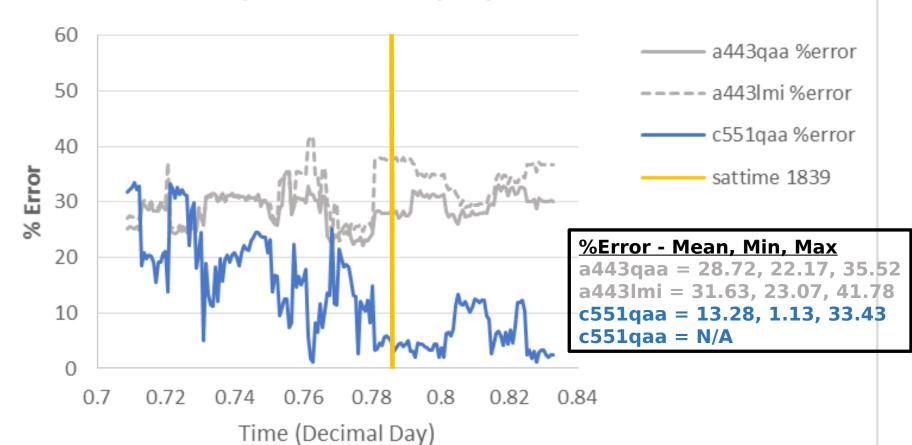


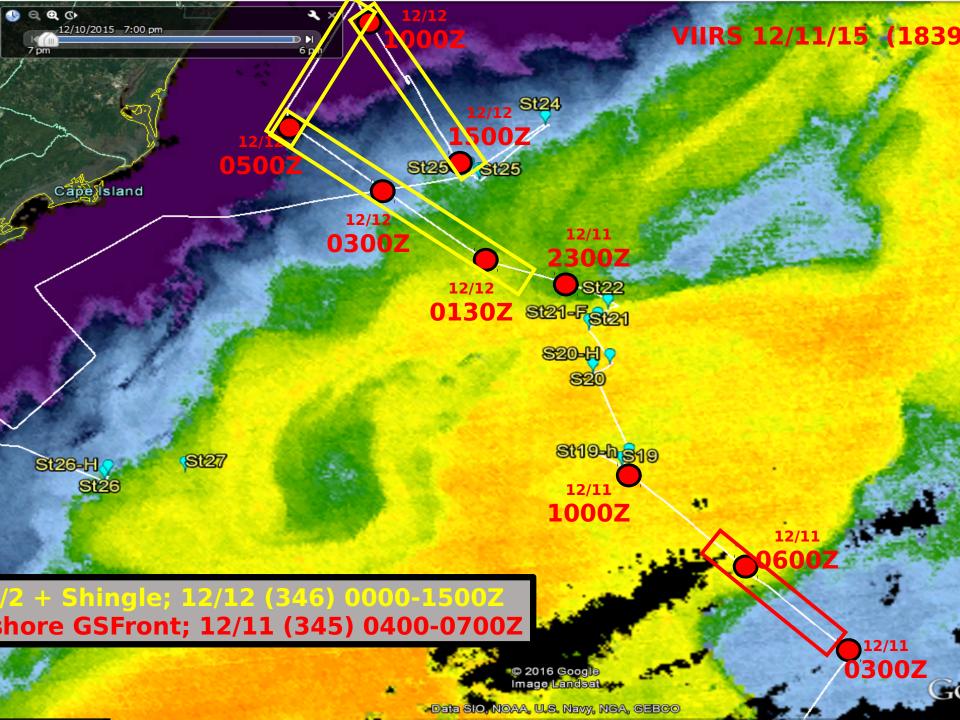


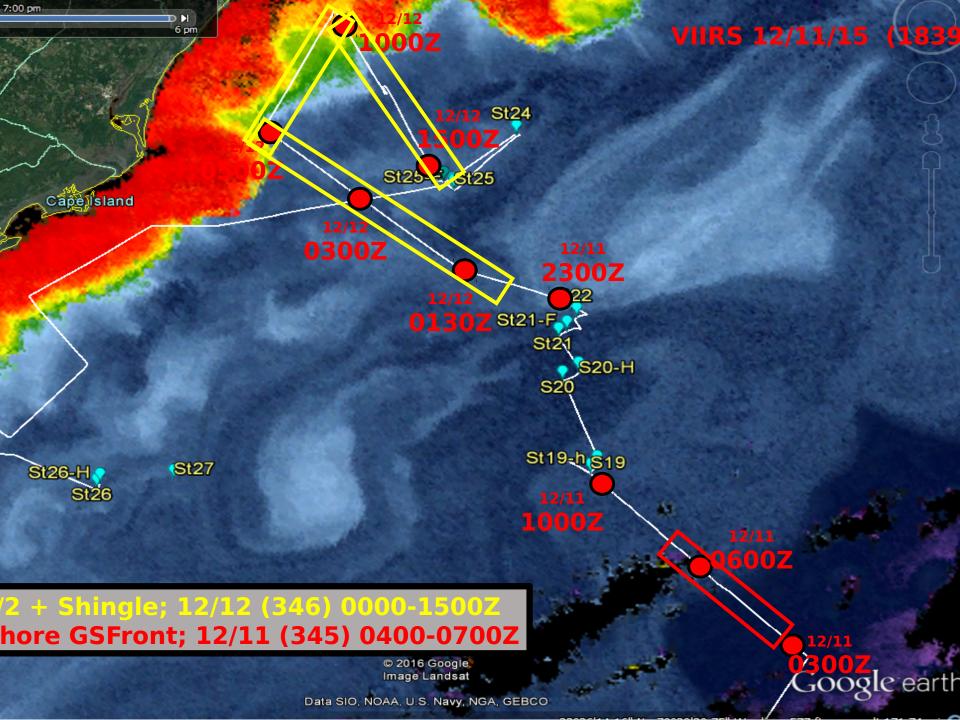


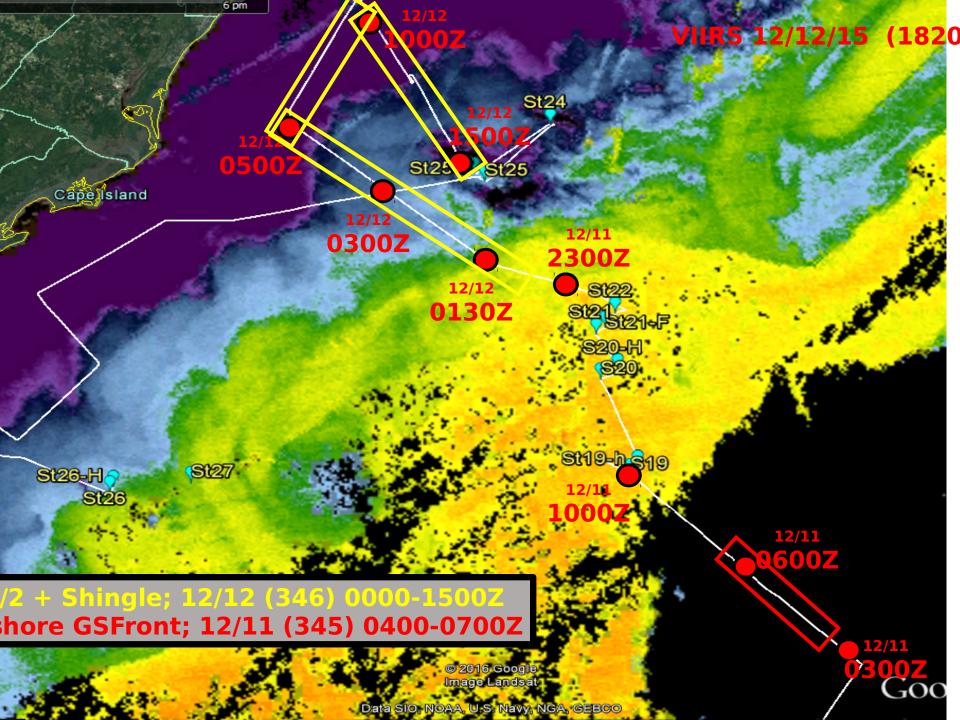












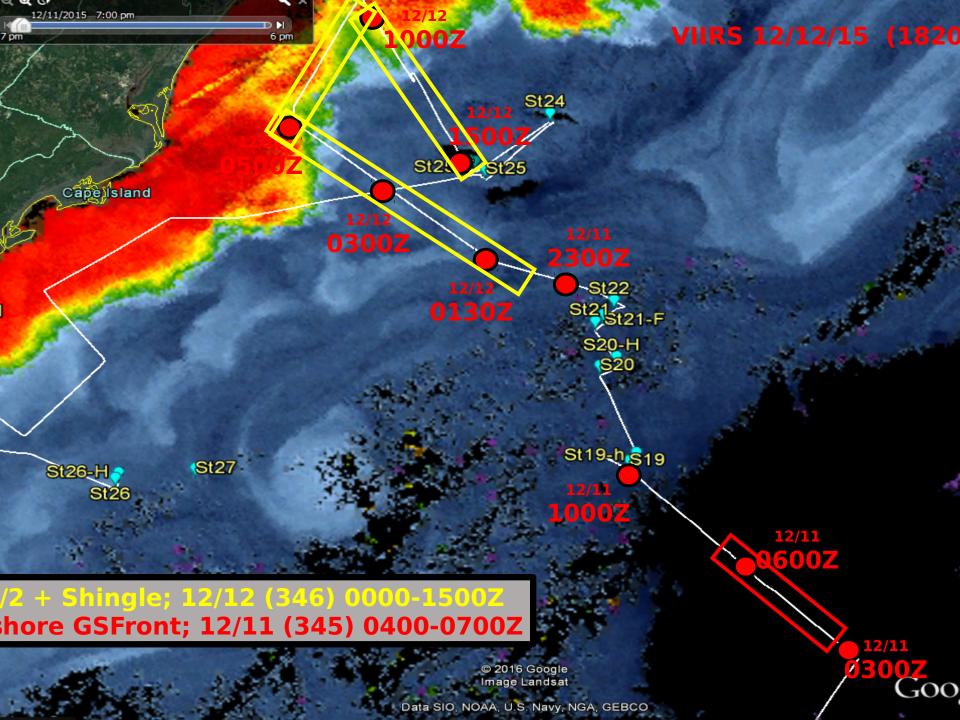
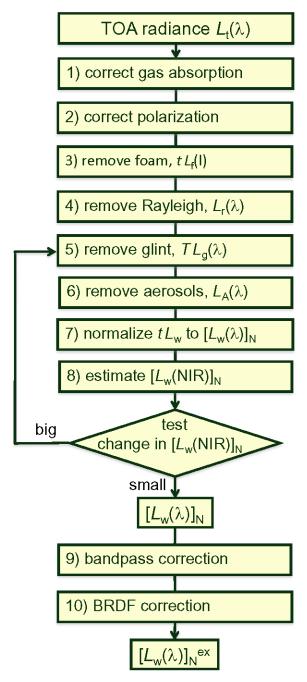
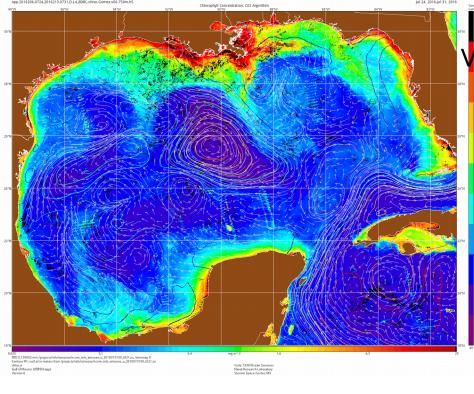


Figure: Flowchart of the atmospheric correction process.

It is important to keep in mind that there are severe computational constraints on how atmospheric correction is performed on an operational basis. The MODIS-Agua sensor, for example, collects about 1.4 terrabytes of data per day. The requirement to routinely process this amount of data (along with data from other sensors) requires that various approximations be made in order to speed up the calculations. Some of the corrections require ancillary information such as sea level pressure, wind speed, and ozone concentration, which are not collected by ocean color sensors themselves. These ancillary data may be inaccurate or missing, in which case climatological values must be used. The quality of the ancillary information impacts the accuracy of the atmospheric correction. Table [*] shows some of the ancillary data and its sources as used by the various OBPG atmospheric correction algorithms.





MODIS Vicariously Calibrated to MOBY
IIRS Vicariously Calibrated to MOBY+AO
Spring 2016 Gain Sets
Inter-sensor Consistency

VIIRS 8 Day Rolling Composite for July 31, 2016
Chlorophyll

MODIS 8 Day Rolling Composite for July 31, 2016 Chlorophyll

